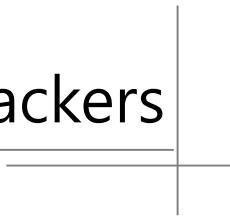
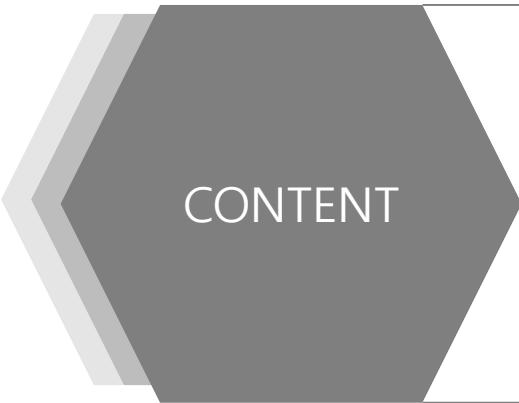




Star-Image Centroiding And Attitude Determination for Star Trackers





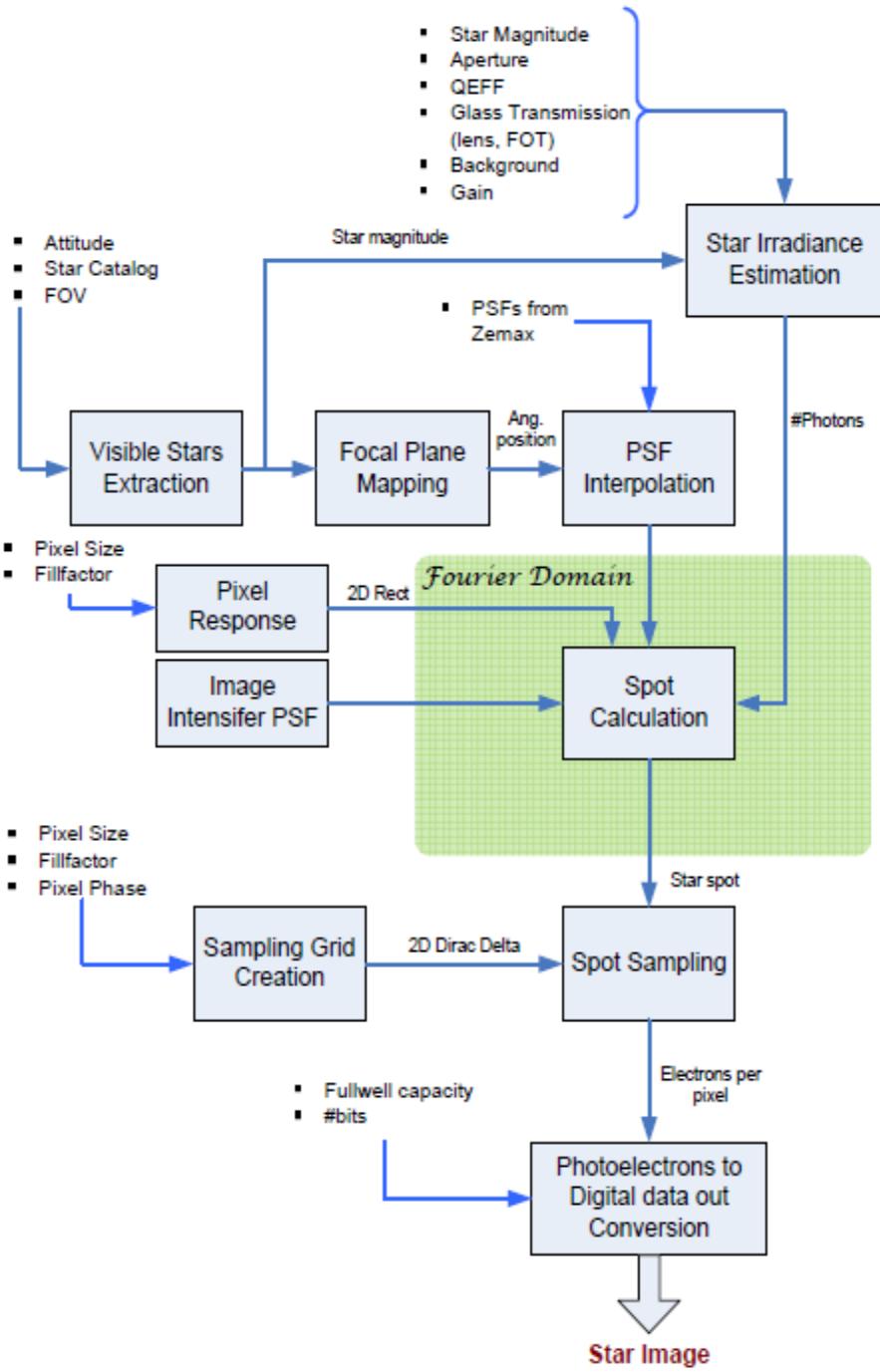
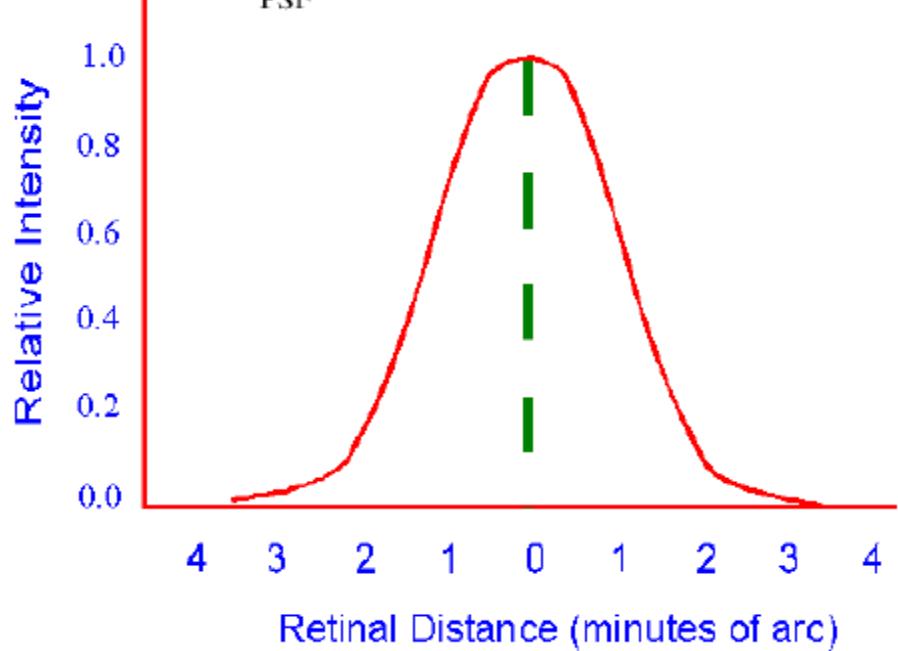
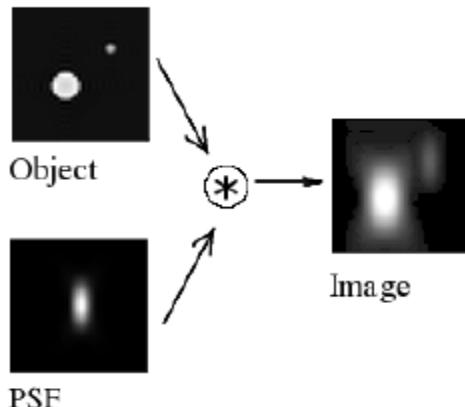
CONTENT

Electro – Optical Modeling And Image Simulator Design

Star Position Estimation Improvements

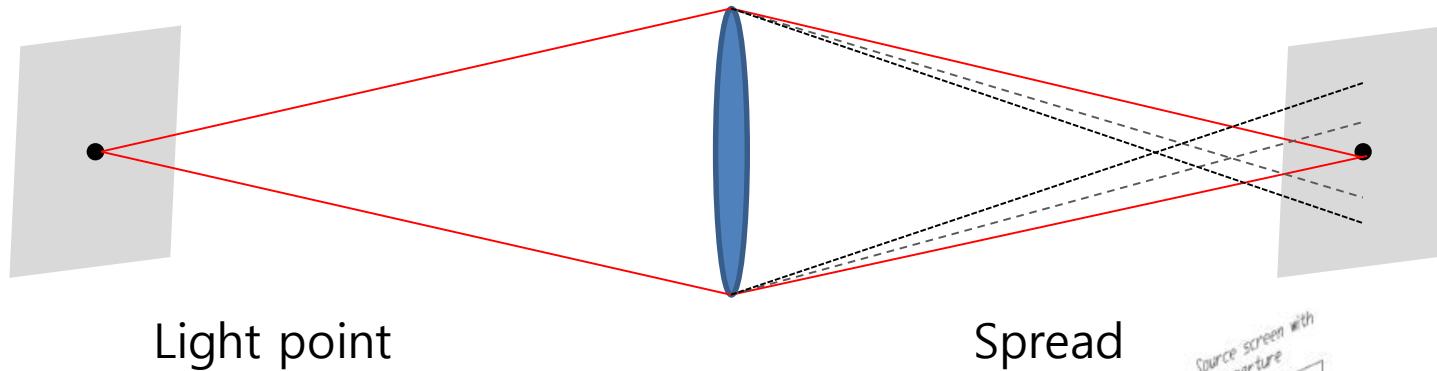
Simulation

Electro – Optical Modeling And Image Simulator Design

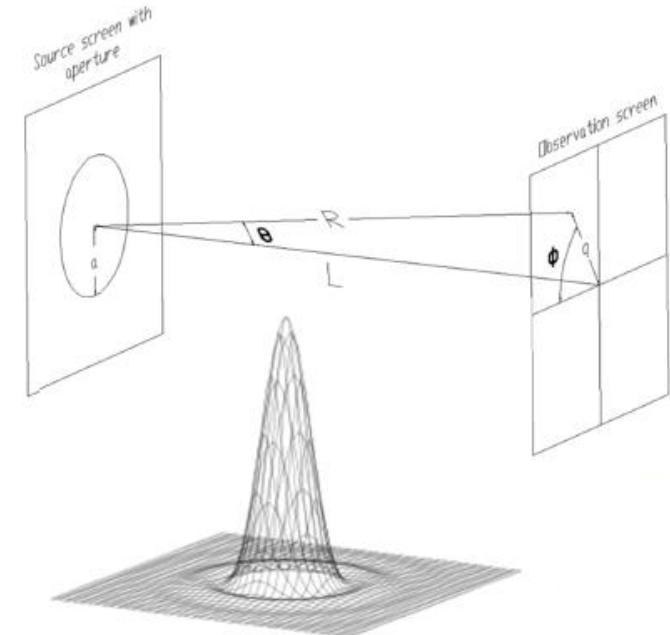


Electro – Optical Modeling And Image Simulator Design

Point Spread Function (PSF)



-빛의 확산인 빛의 분포 정도를 Point Spread Function(PSF)라고 부른다.



Electro – Optical Modeling And Image Simulator Design

- 2D Gaussian function

$$f(x, y) = \frac{E}{(2\pi\sigma_x\sigma_y)} e^{-\frac{(x-x_0)^2}{2\sigma_x^2}} e^{-\frac{(y-y_0)^2}{2\sigma_y^2}}$$

where E , $\square t \times e^{\frac{(m_0-m)}{2.5}}$ \longrightarrow Star energy

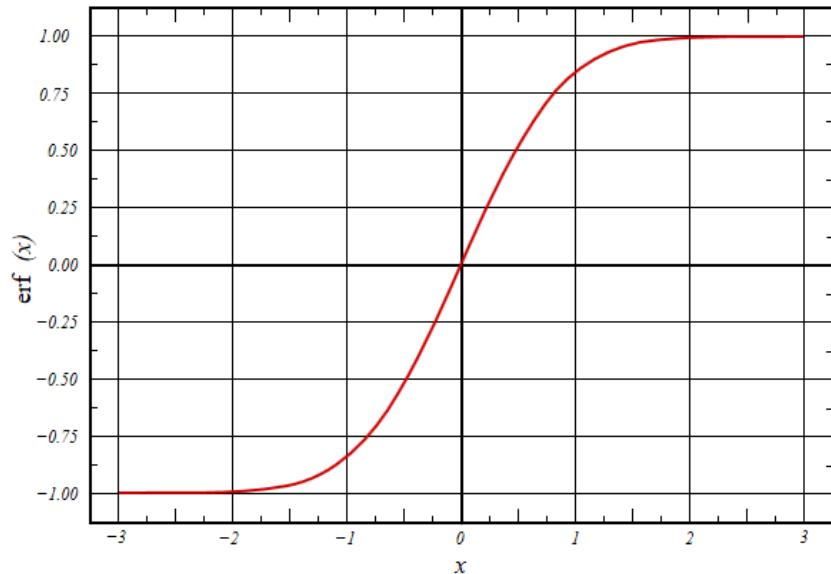
- Error function (ERF)

Electro – Optical Modeling And Image Simulator Design

- Error function (ERF)

$$\text{erf}(x) = \frac{2}{\sqrt{\pi}} \int_0^x e^{-t^2} dt$$

→ $\int_0^\infty e^{-\frac{(x-x_0)^2}{2\sigma_x^2}} dt = \frac{2}{\sqrt{\pi}} \left[1 - \text{erf}\left(\frac{x_0 - h}{\sigma_x \sqrt{2}}\right) \right] \sigma_x$



Electro – Optical Modeling And Image Simulator Design

$$f(x, y) = \frac{E}{(2\pi\sigma_x\sigma_y)} e^{-\frac{(x-x_0)^2}{2\sigma_x^2}} e^{-\frac{(y-y_0)^2}{2\sigma_y^2}}$$

$$\int_{x_1}^{x_2} e^{-\frac{(x-x_0)^2}{2\sigma_x^2}} dx = \int_{x_1}^{\infty} e^{-\frac{(x-x_0)^2}{2\sigma_x^2}} dx - \int_{x_2}^{\infty} e^{-\frac{(x-x_0)^2}{2\sigma_x^2}} dx$$

$$= \sqrt{\frac{\pi}{2}} [\varepsilon(x_2) - \varepsilon(x_1)] \sigma_x$$

where $\varepsilon(x_1)$, $\operatorname{erf}\left(\frac{x_0 - x_1}{\sigma_x \sqrt{2}}\right)$ $\varepsilon(x_2)$, $\operatorname{erf}\left(\frac{x_0 - x_2}{\sigma_x \sqrt{2}}\right)$

Electro – Optical Modeling And Image Simulator Design

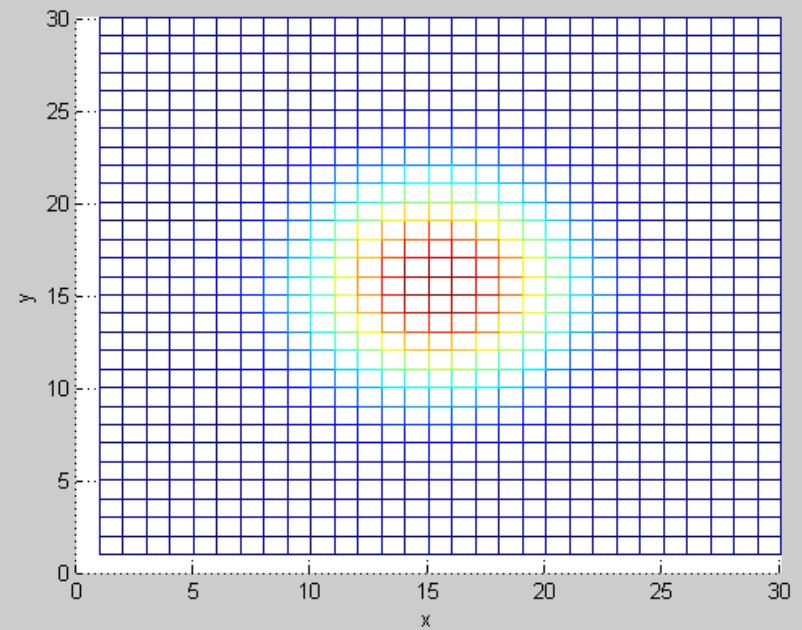
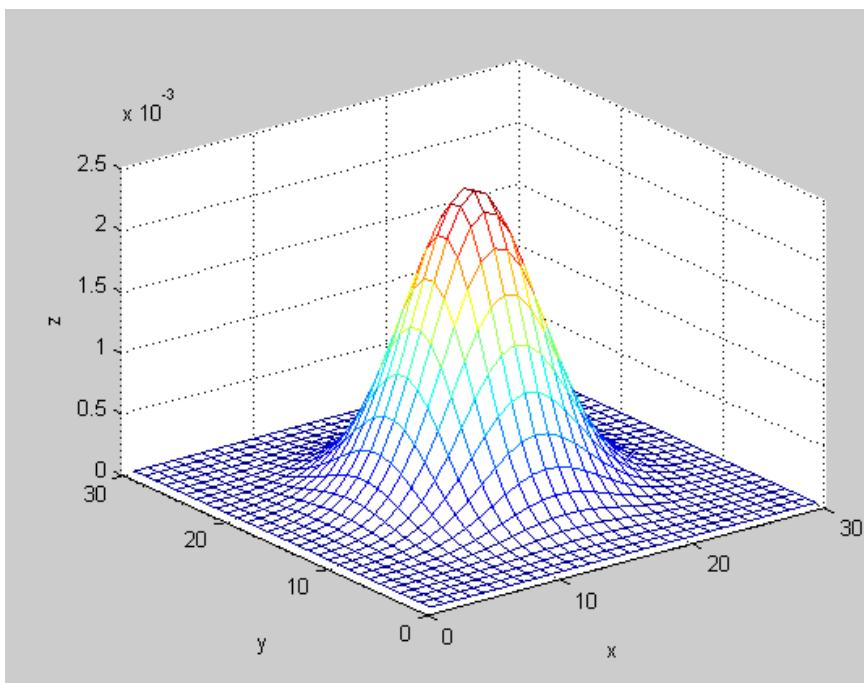
$$f(x, y) = \frac{E}{(2\pi\sigma_x\sigma_y)} e^{-\frac{(x-x_0)^2}{2\sigma_x^2}} e^{-\frac{(y-y_0)^2}{2\sigma_y^2}}$$

$$\Rightarrow \frac{E}{(2\pi\sigma_x\sigma_y)} \left[\int_{y_1}^{y_2} e^{-\frac{(y-y_0)^2}{2\sigma_y^2}} dy \right] \left[\int_{x_1}^{x_2} e^{-\frac{(x-x_0)^2}{2\sigma_x^2}} dx \right]$$

$$\Rightarrow \frac{E}{4} [\varepsilon(y_2) - \varepsilon(y_1)] [\varepsilon(x_2) - \varepsilon(x_1)]$$

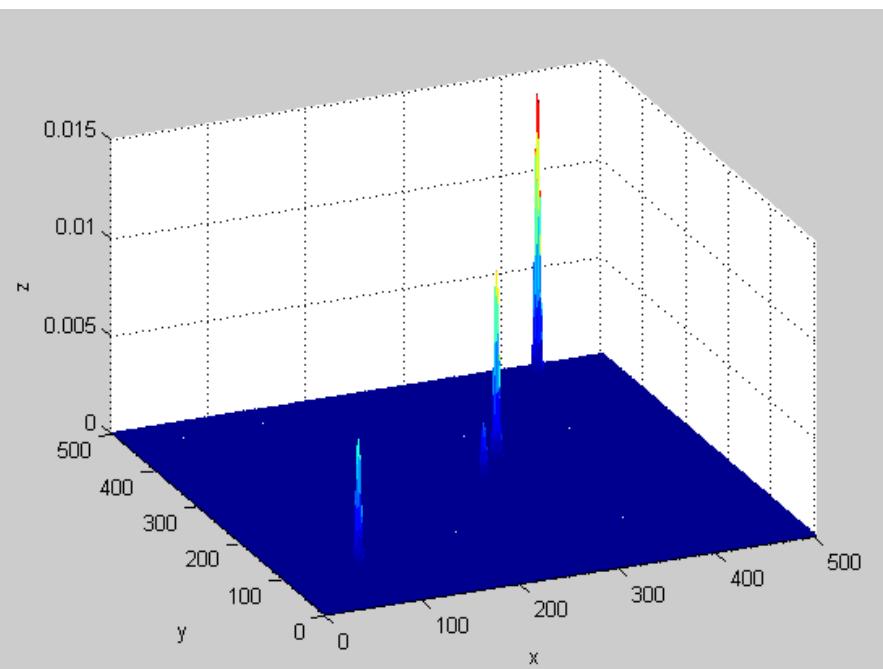
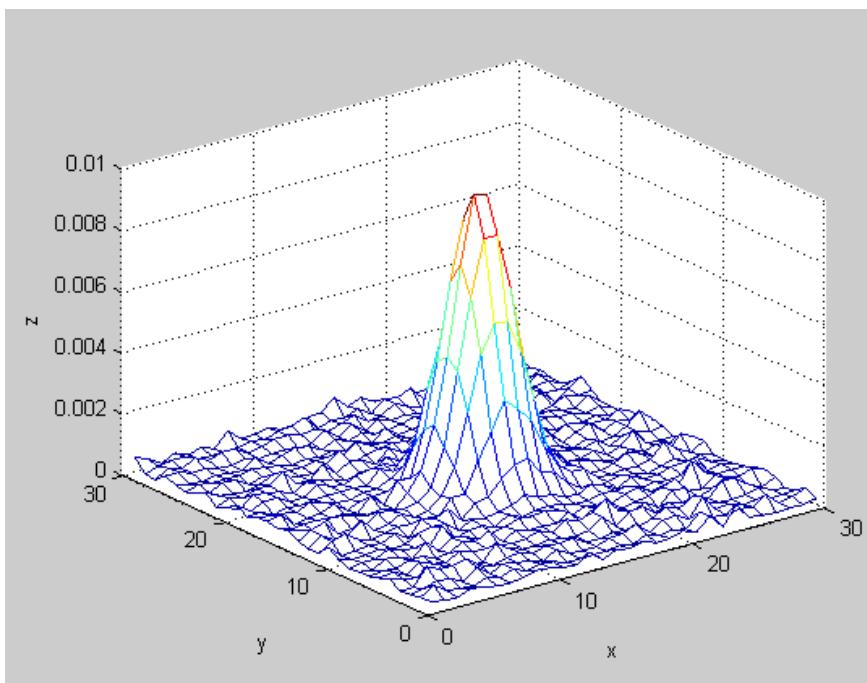
Electro – Optical Modeling And Image Simulator Design

- Simulation



Electro – Optical Modeling And Image Simulator Design

- Simulation



Star Position Estimation Improvements

- Methods

Center
Of
Mass

Non-Linear
Gaussian
Fitting

Linear Least
Squares
Log of
Gaussian
Estimation

$$x_{CM} = \frac{\sum_{i,j} x_{ij} E_{ij}}{\sum_{i,j} E_{ij}}$$

$$\Delta_{k+1} = [F_k^T F_k]^{-1} F_k^T R_k$$

$$z = (A^T W A)^{-1} A^T W F$$

Star Position Estimation Improvements

- Center Of Mass(COM)

$$x_c = x_M + \frac{\sum_{i=-n/2}^{n/2} \sum_{j=-m/2}^{m/2} [P(x_M - i, y_M - j) - B] i}{\sum_{i=-n/2}^{n/2} \sum_{j=-m/2}^{m/2} [P(x_M - i, y_M - j) - B]}$$

$$y_c = y_M + \frac{\sum_{i=-n/2}^{n/2} \sum_{j=-m/2}^{m/2} [P(x_M - i, y_M - j) - B] j}{\sum_{i=-n/2}^{n/2} \sum_{j=-m/2}^{m/2} [P(x_M - i, y_M - j) - B]}$$

$(x_M \ y_M)$: location of the peak of the intensity distribution

n, m : mask sizes along x and y directions

B : Background noise intensity (constant)

Star Position Estimation Improvements

- Non-Linear 2-D Gaussian Fitting

$$P(x, y) = \frac{t_{int}}{2\pi\sigma_x\sigma_y} e^{-\left(\frac{m_0 - m}{2.5}\right)^2} e^{-\frac{(x - x_c)^2}{2\sigma_x^2}} e^{-\frac{(y - y_c)^2}{2\sigma_y^2}}$$

$$\Rightarrow \ln P(x, y) = \underbrace{\ln \frac{t_{int}}{2\pi\sigma_x\sigma_y}}_{z_c} - \frac{m_0 - m}{2.5} - \frac{(x - x_c)^2}{2\sigma_x^2} - \frac{(y - y_c)^2}{2\sigma_y^2}$$

$$\Rightarrow z_i = z_c - \frac{(x_i - x_c)^2}{2\sigma_x^2} - \frac{(y_i - y_c)^2}{2\sigma_y^2} = f(x_c, y_c, z_c)$$

$$\epsilon_i = z_i - f(x_c, y_c, z_c) \quad L = \sum_{i=1}^n \alpha_i \epsilon_i^2 \quad \text{where } \alpha, \text{relative weights}$$

Star Position Estimation Improvements

- Non-Linear 2-D Gaussian Fitting

$$\frac{\partial L}{\partial x_0} = 0, \quad \frac{\partial L}{\partial y_0} = 0, \quad \frac{\partial L}{\partial z_0} = 0$$

$$\frac{\partial L}{\partial x_c} = 0 \Rightarrow \sum_i [2(z_i - z_c) + \sigma_x^{-2} (x_i - x_c)^2 + \sigma_y^{-2} (y_i - y_c)^2] (x_i - x_c) = 0$$

→ $\frac{\partial L}{\partial y_c} = 0 \Rightarrow \sum_i [2(z_i - z_c) + \sigma_x^{-2} (x_i - x_c)^2 + \sigma_y^{-2} (y_i - y_c)^2] (y_i - y_c) = 0$

$$\frac{\partial L}{\partial z_c} = 0 \Rightarrow \sum_i [2(z_i - z_c) + \sigma_x^{-2} (x_i - x_c)^2 + \sigma_y^{-2} (y_i - y_c)^2] = 0$$

$$r_i = z_i - f_i(\hat{x}_0, \hat{y}_0, \hat{z}_0) = \begin{Bmatrix} \frac{\partial f_i}{\partial x_0} & \frac{\partial f_i}{\partial y_0} & \frac{\partial f_i}{\partial z_0} \end{Bmatrix} \begin{Bmatrix} \delta x_0 \\ \delta y_0 \\ \delta z_0 \end{Bmatrix} = F_i^T \Delta$$

Star Position Estimation Improvements

- Non-Linear 2-D Gaussian Fitting

$$R = \begin{Bmatrix} r_1 \\ \vdots \\ r_n \end{Bmatrix} = \begin{bmatrix} f_{1x} & f_{1y} & f_{1z} \\ \vdots & \vdots & \vdots \\ f_{nx} & f_{ny} & f_{nz} \end{bmatrix} \begin{Bmatrix} \delta x_0 \\ \delta y_0 \\ \delta z_0 \end{Bmatrix} = F\Delta$$

$$\Delta_{k+1} = [F_k^T F_k]^{-1} F_k^T R_k$$

Star Position Estimation Improvements

- Linear Least Squares Log of Gaussian Estimation(LOG-LSQ)

$\mathbf{X} = [x \ y]^T$: pixel location

$$f(\mathbf{X}, \mathbf{X}_c) = \frac{1}{2\pi\sqrt{\det R}} e^{-\frac{1}{2}(\mathbf{X}-\mathbf{X}_c)^T R^{-1} (\mathbf{X}-\mathbf{X}_c)}$$

$\mathbf{X}_c = [x_c \ y_c]^T$: pixel location of the centroid

$R : 2 \times 2$ symmetric covariance matrix

$$\Rightarrow -\ln f = -\ln \frac{1}{2\pi\sqrt{\det R}} + \frac{1}{2} (\mathbf{X} - \mathbf{X}_c)^T R^{-1} (\mathbf{X} - \mathbf{X}_c)$$

$$-\ln f = -\underbrace{\ln \frac{1}{2\pi\sqrt{\det R}}}_{a_0} + \frac{1}{2} \mathbf{X}_c^T D \mathbf{X}_c + \frac{1}{2} \mathbf{X}^T D \mathbf{X} - \mathbf{X}_c^T D \mathbf{X}$$

where $D = R^{-1}$

$$-\ln f = a_0 + \frac{1}{2} \mathbf{X}^T D \mathbf{X} - \mathbf{X}_c^T D \mathbf{X}$$

Star Position Estimation Improvements

- Linear Least Squares Log of Gaussian Estimation(LOG-LSQ)

$$-\ln f = a_0 + \begin{bmatrix} a_1 & a_2 \end{bmatrix} \begin{bmatrix} x \\ y \end{bmatrix} + \frac{1}{2} \begin{bmatrix} x & y \end{bmatrix} \begin{bmatrix} d_{11} & d_{12} \\ d_{21} & d_{22} \end{bmatrix} \begin{bmatrix} x \\ y \end{bmatrix}$$

where $[a_1 \ a_2] = \mathbf{X}_c^T D$

$$-\ln f = a_0 - a_1x - a_2y + \frac{1}{2}d_{11}x^2 + \frac{1}{2}d_{12}xy + \frac{1}{2}d_{21}xy + \frac{1}{2}d_{22}y^2$$

$$\begin{bmatrix} -\ln f_1 \\ -\ln f_2 \\ \vdots \\ -\ln f_n \end{bmatrix} = \begin{bmatrix} 1 & -x_1 & -y_1 & \frac{1}{2}x_1^2 & \frac{1}{2}x_1y_1 & \frac{1}{2}x_1y_1 & \frac{1}{2}y_1^2 \\ 1 & -x_2 & -y_2 & \frac{1}{2}x_2^2 & \frac{1}{2}x_2y_2 & \frac{1}{2}x_2y_2 & \frac{1}{2}y_2^2 \\ \vdots & \vdots & \vdots & \vdots & \vdots & \vdots & \vdots \\ 1 & -x_n & -y_n & \frac{1}{2}x_n^2 & \frac{1}{2}x_ny_n & \frac{1}{2}x_ny_n & \frac{1}{2}y_n^2 \end{bmatrix} \begin{bmatrix} a_0 \\ a_1 \\ a_2 \\ d_{11} \\ d_{12} \\ d_{21} \\ d_{22} \end{bmatrix}$$

$$\tilde{f} = A'z'$$

Star Position Estimation Improvements

- Linear Least Squares Log of Gaussian Estimation(LOG-LSQ)

$$R = \begin{bmatrix} \sigma_x^2 & \rho\sigma_x\sigma_y \\ \rho\sigma_x\sigma_y & \sigma_y^2 \end{bmatrix}$$

where ρ , The correlation between the x and y directions

$$\begin{bmatrix} -\ln f_1 \\ -\ln f_2 \\ \vdots \\ -\ln f_n \end{bmatrix} = \begin{bmatrix} 1 & -x_1 & -y_1 & \frac{1}{2}x_1^2 & x_1y_1 & \frac{1}{2}y_1^2 \\ 1 & -x_2 & -y_2 & \frac{1}{2}x_2^2 & x_2y_2 & \frac{1}{2}y_2^2 \\ \vdots & \vdots & \vdots & \vdots & \vdots & \vdots \\ 1 & -x_n & -y_n & \frac{1}{2}x_n^2 & x_ny_n & \frac{1}{2}y_n^2 \end{bmatrix} \begin{bmatrix} a_0 \\ a_1 \\ a_2 \\ d_{11} \\ d_2 \\ d_{22} \end{bmatrix}$$

$$\tilde{f} = A\mathbf{z}$$

Star Position Estimation Improvements

- Linear Least Squares Log of Gaussian Estimation(LOG-LSQ)

$$\begin{bmatrix} -\ln f_1 \\ -\ln f_2 \\ \vdots \\ -\ln f_n \end{bmatrix} = \begin{bmatrix} 1 & -x_1 & -y_1 & \frac{1}{2}x_1^2 & x_1y_1 & \frac{1}{2}y_1^2 \\ 1 & -x_2 & -y_2 & \frac{1}{2}x_2^2 & x_2y_2 & \frac{1}{2}y_2^2 \\ \vdots & \vdots & \vdots & \vdots & \vdots & \vdots \\ 1 & -x_n & -y_n & \frac{1}{2}x_n^2 & x_ny_n & \frac{1}{2}y_n^2 \end{bmatrix} \begin{bmatrix} a_0 \\ a_1 \\ a_2 \\ d_{11} \\ d_2 \\ d_{22} \end{bmatrix}$$

$$\tilde{f} = Az$$

$$\mathbf{z} = (A^T A)^{-1} A^T F$$

$$X_c^T D = [a_1 \ a_2]$$

$$\Rightarrow \mathbf{X}_c = ([a_1 \ a_2] D^{-1})^T$$

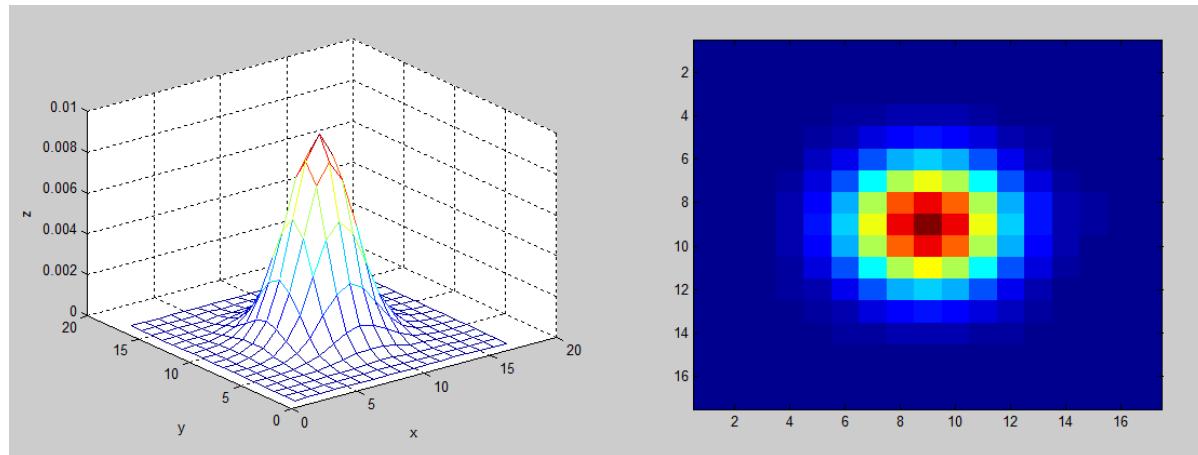
Simulation

- Variable
 - Sigma
 - Magnitude
 - Noise
 - Mask size
 - Weight
 - threshold

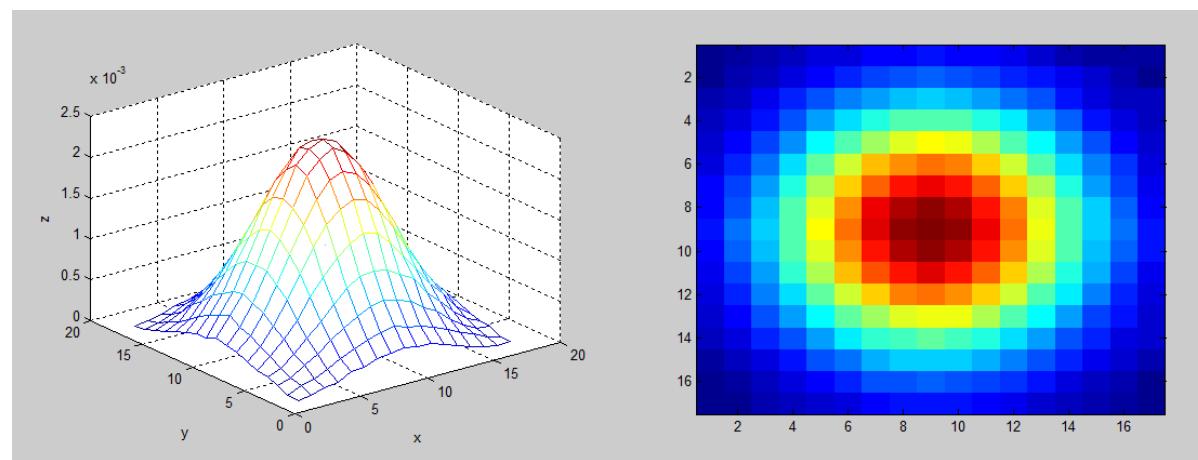
Simulation

- Sigma

Sigma 2

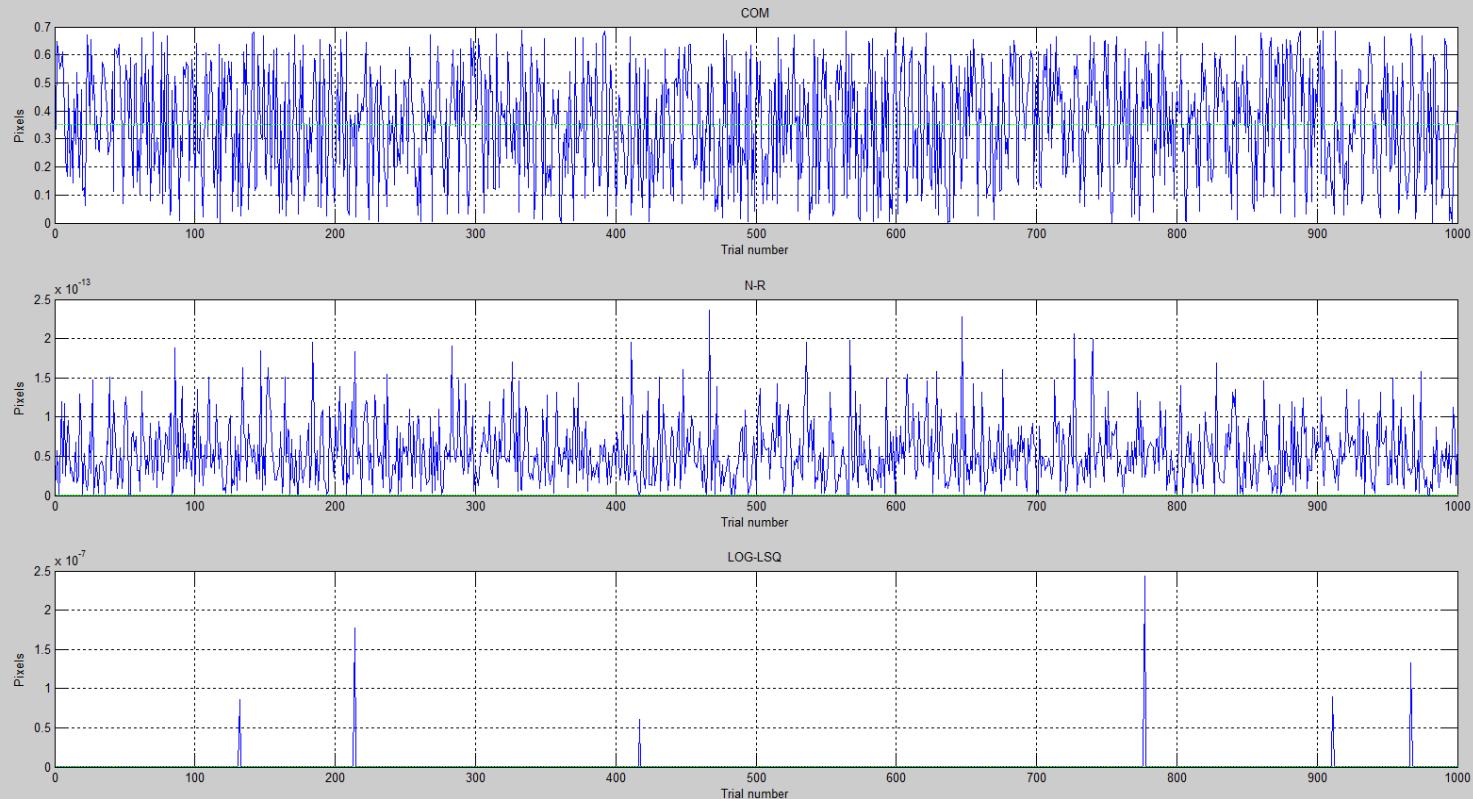


Sigma 4



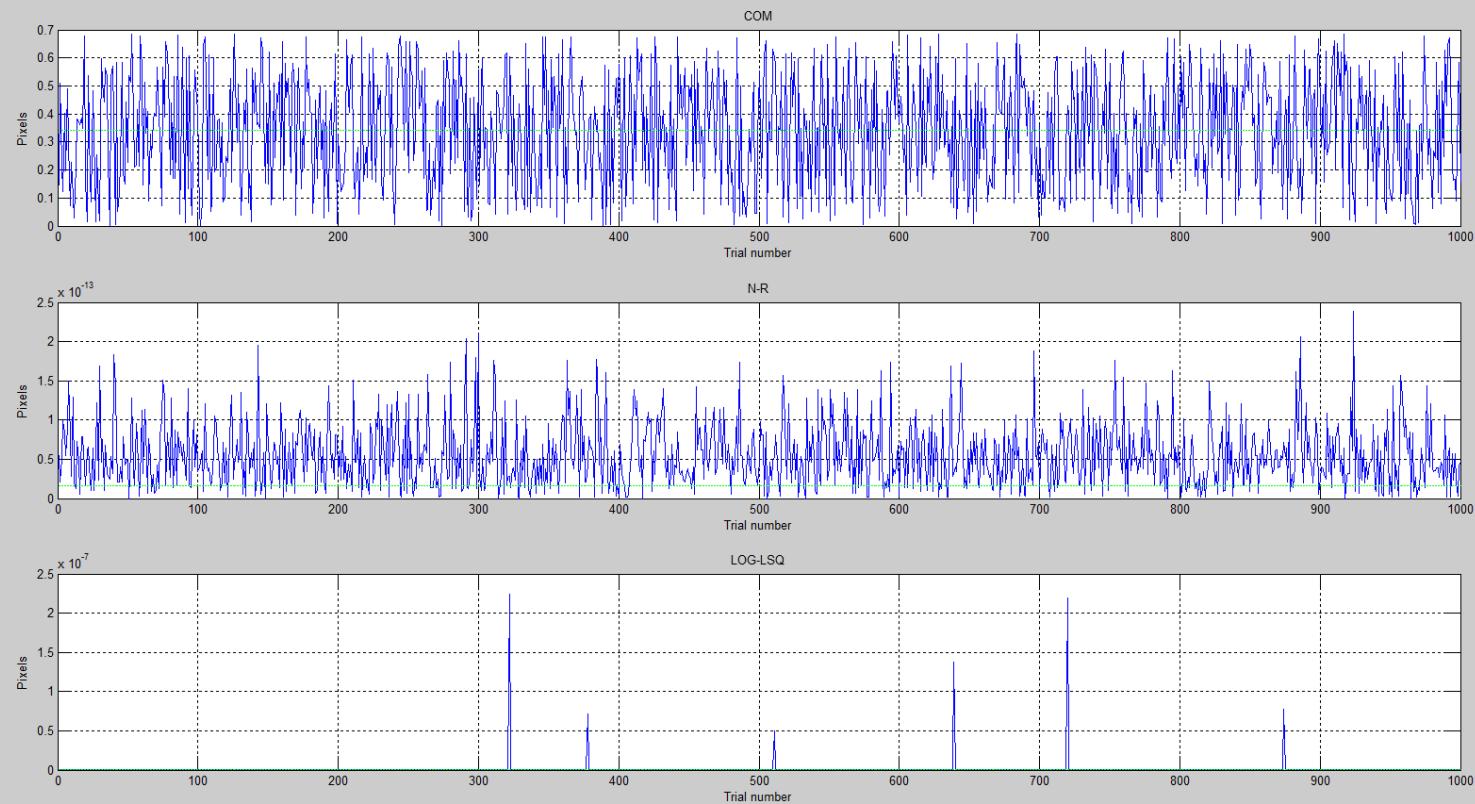
Simulation

- Magnitude (Noise 0%, m -1, Mask 7)



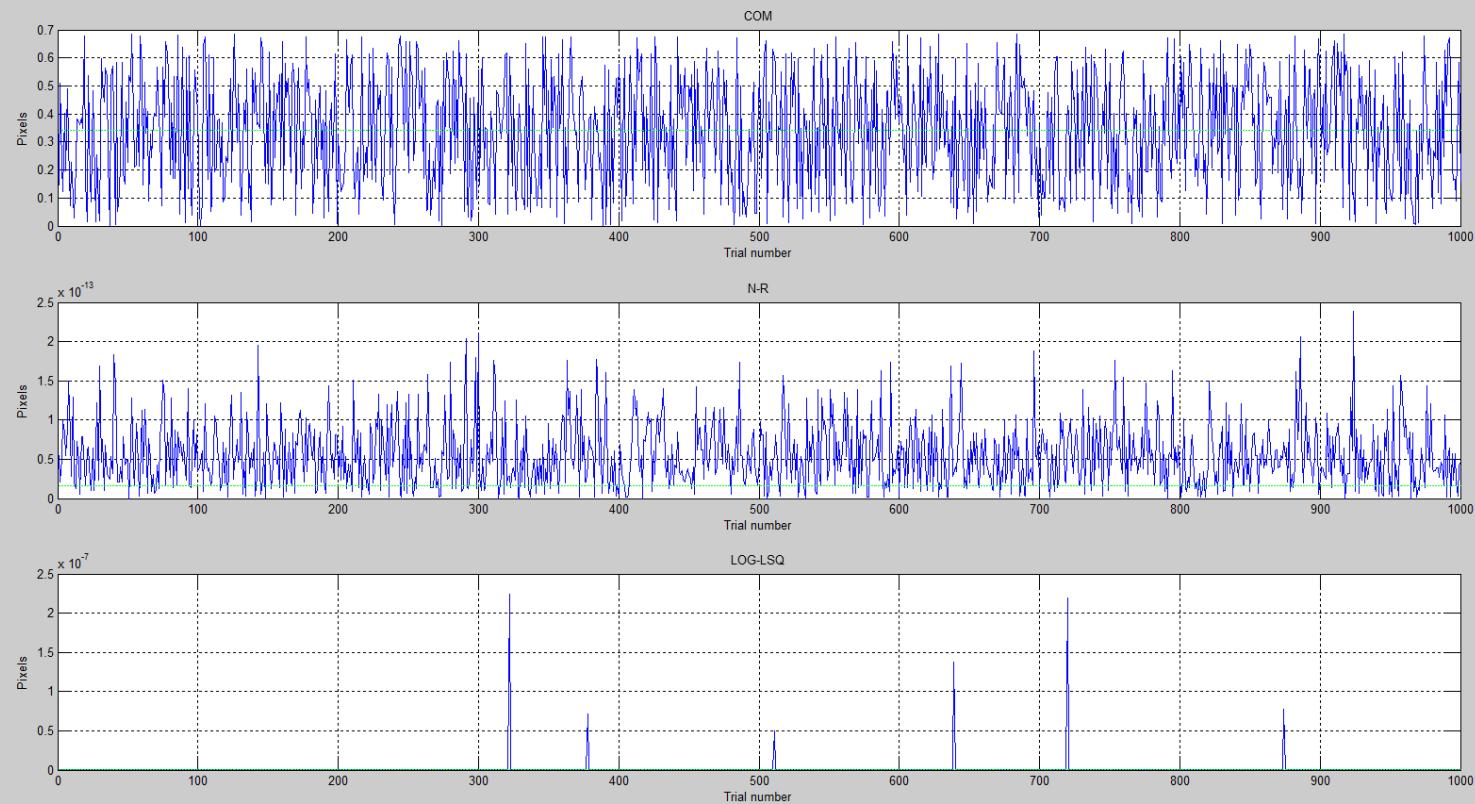
Simulation

- Magnitude (Noise 0%, m 3, Mask 7)



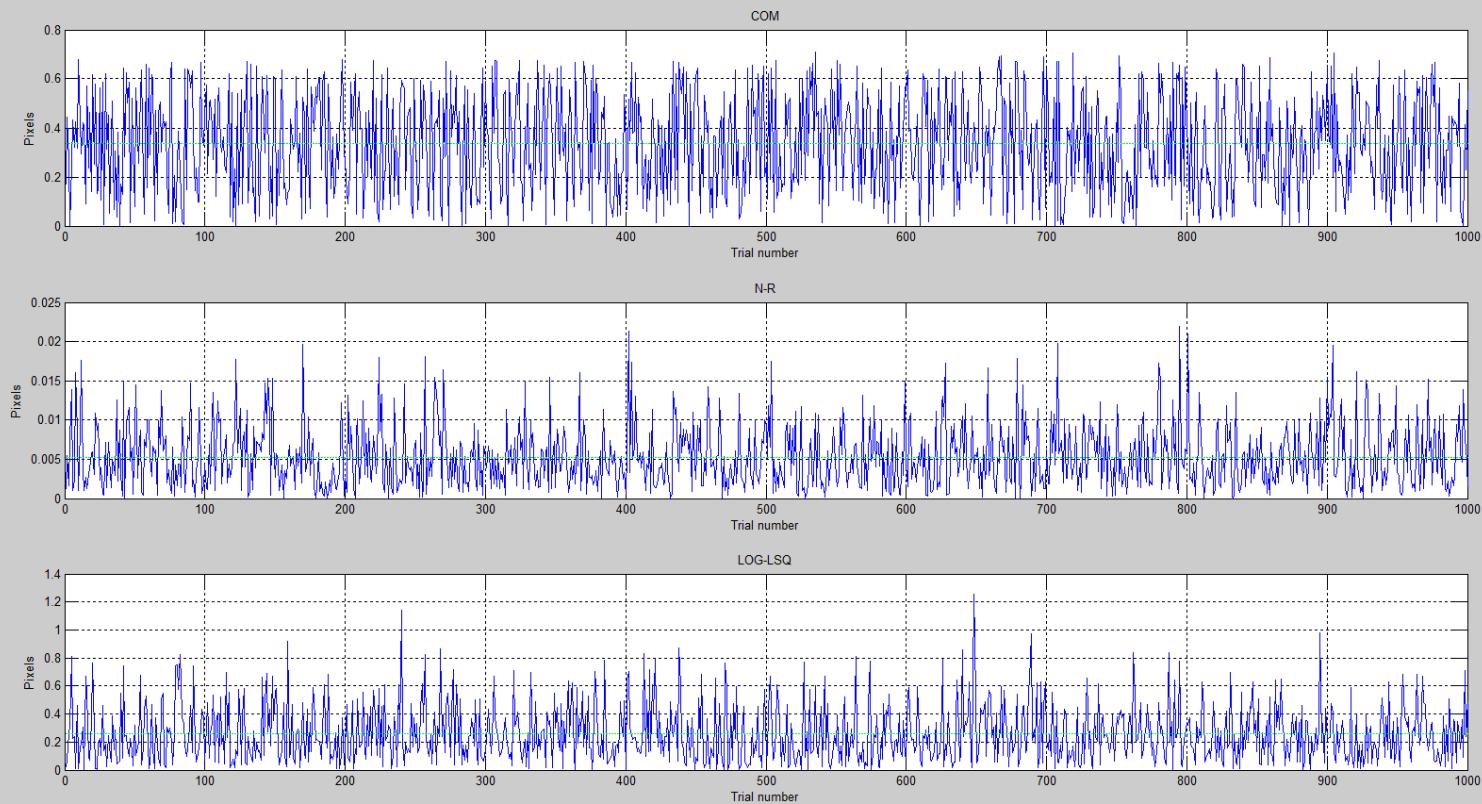
Simulation

- Noise (Noise 0%, m 3, Mask 7)



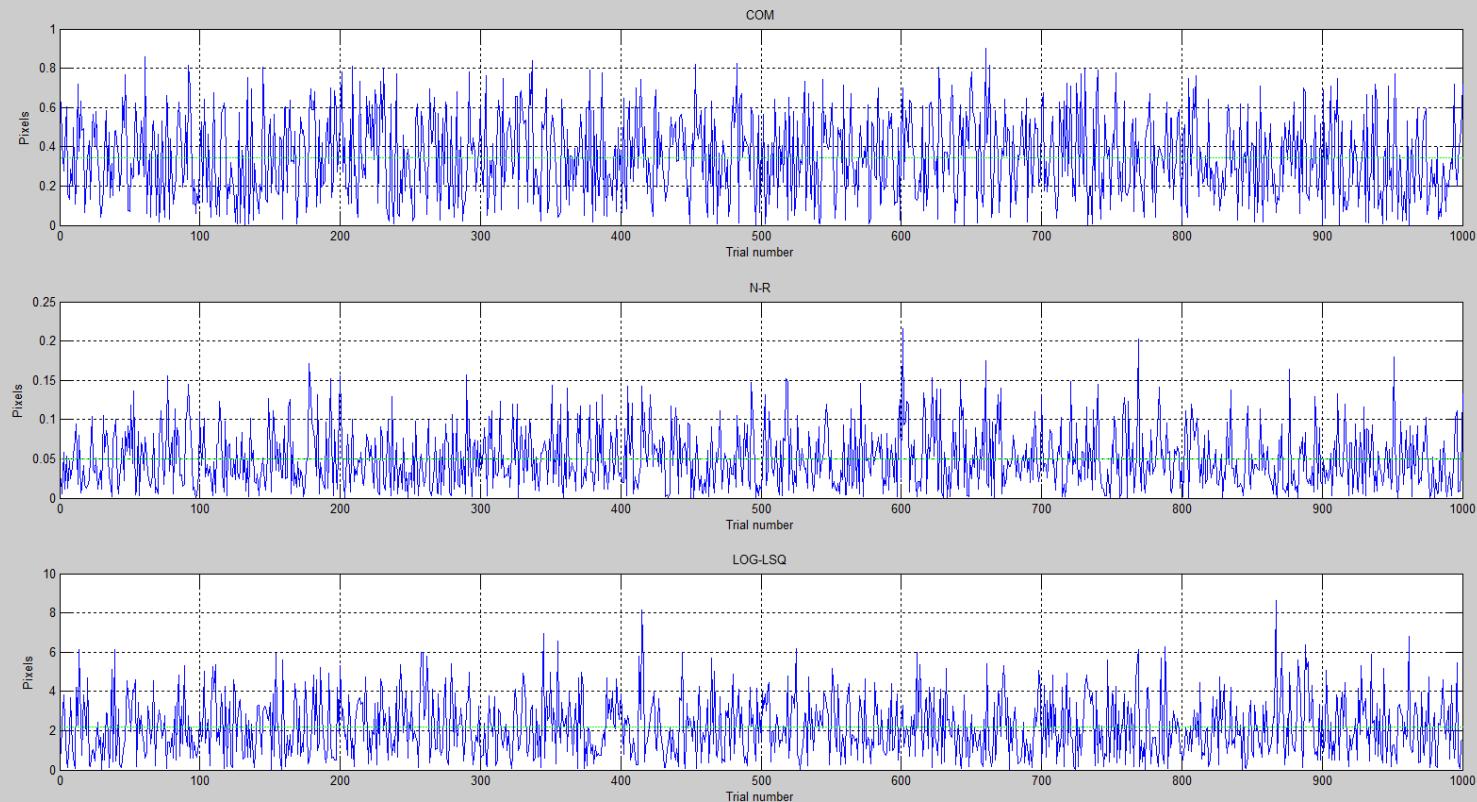
Simulation

- Noise (Noise 1%, m 3, Mask 7)



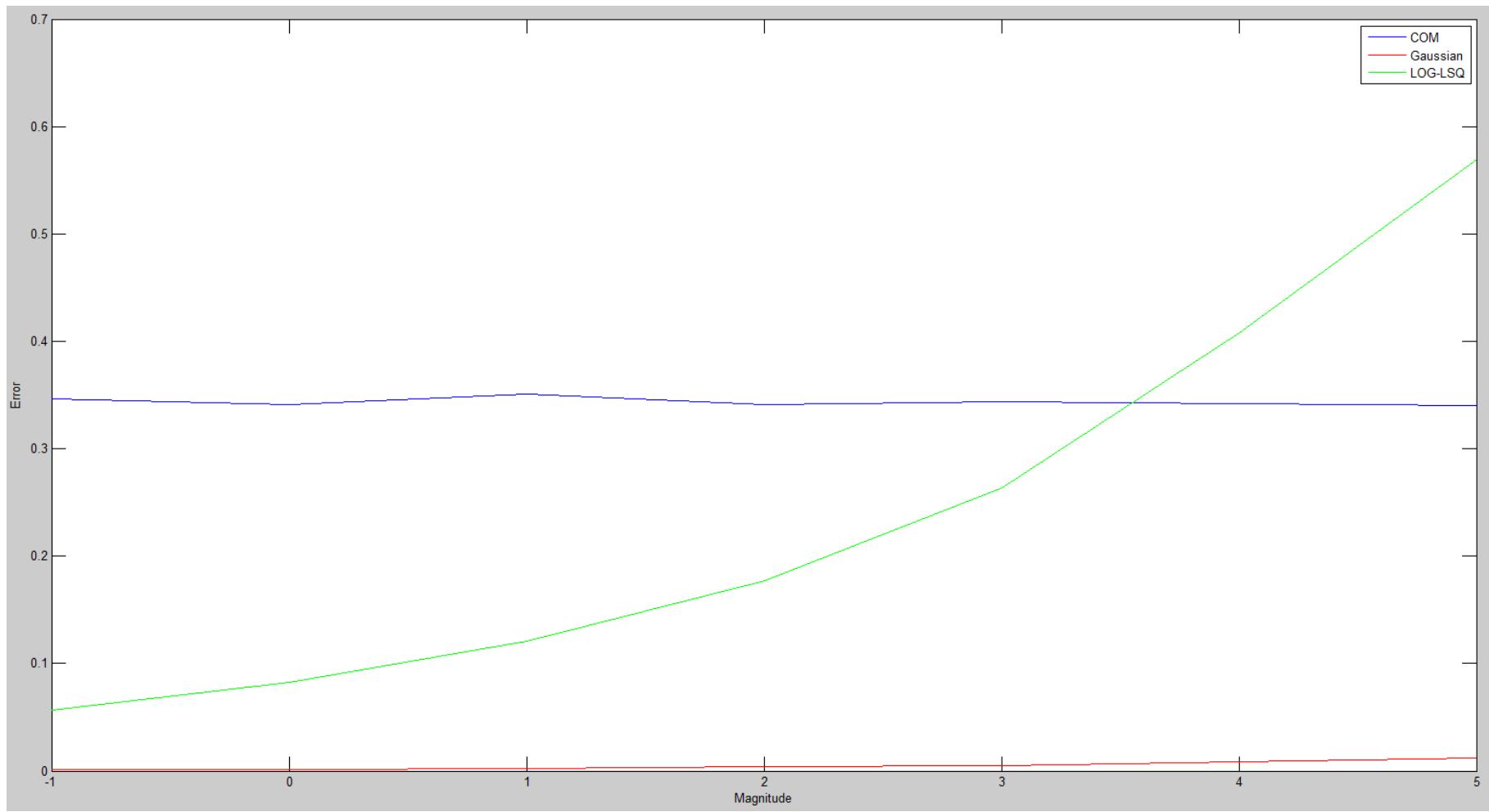
Simulation

- Noise (Noise 10%, m 3, Mask 7)



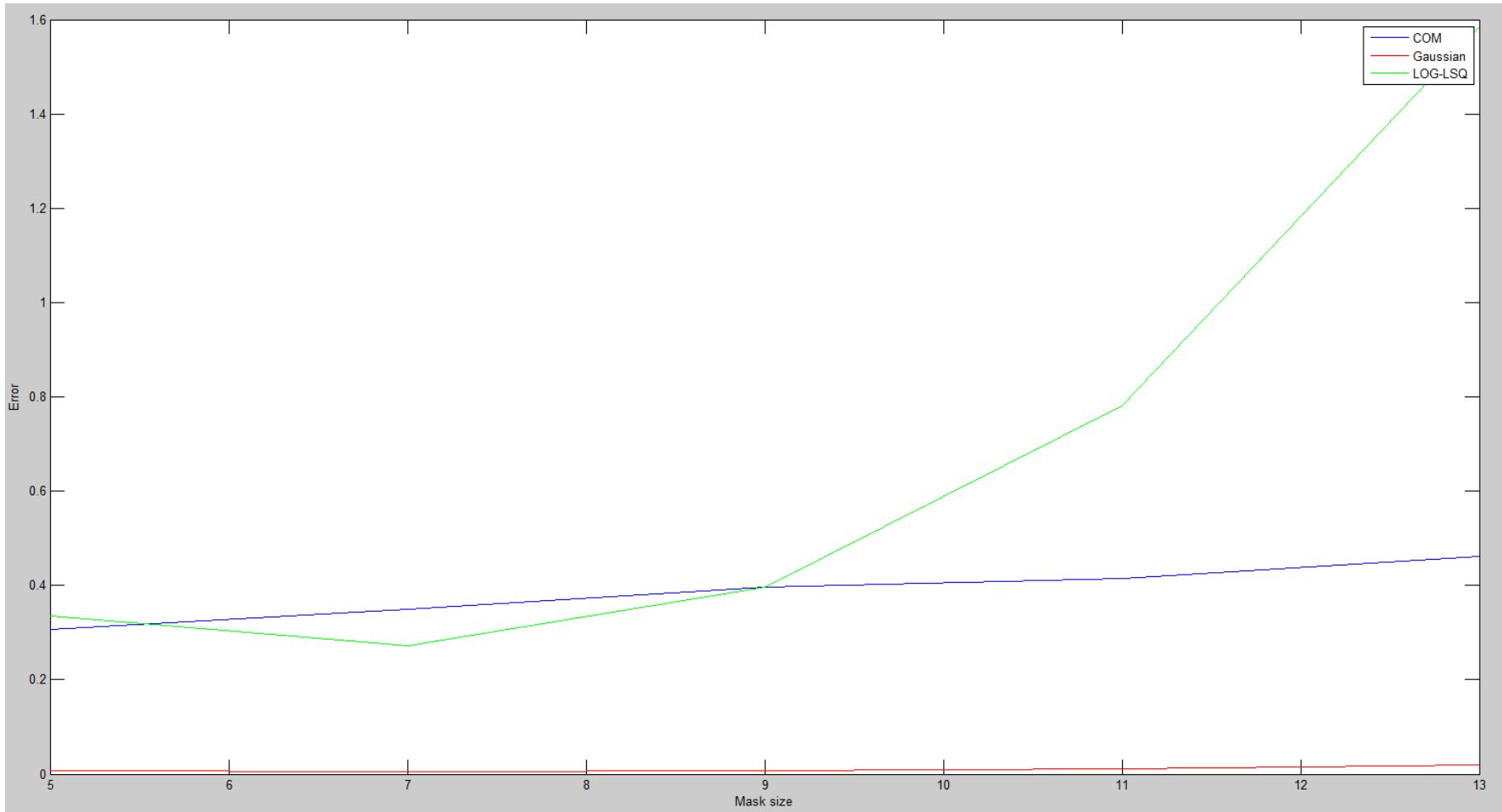
Simulation

- Magnitude Error (Noise 1%, Mask 7)



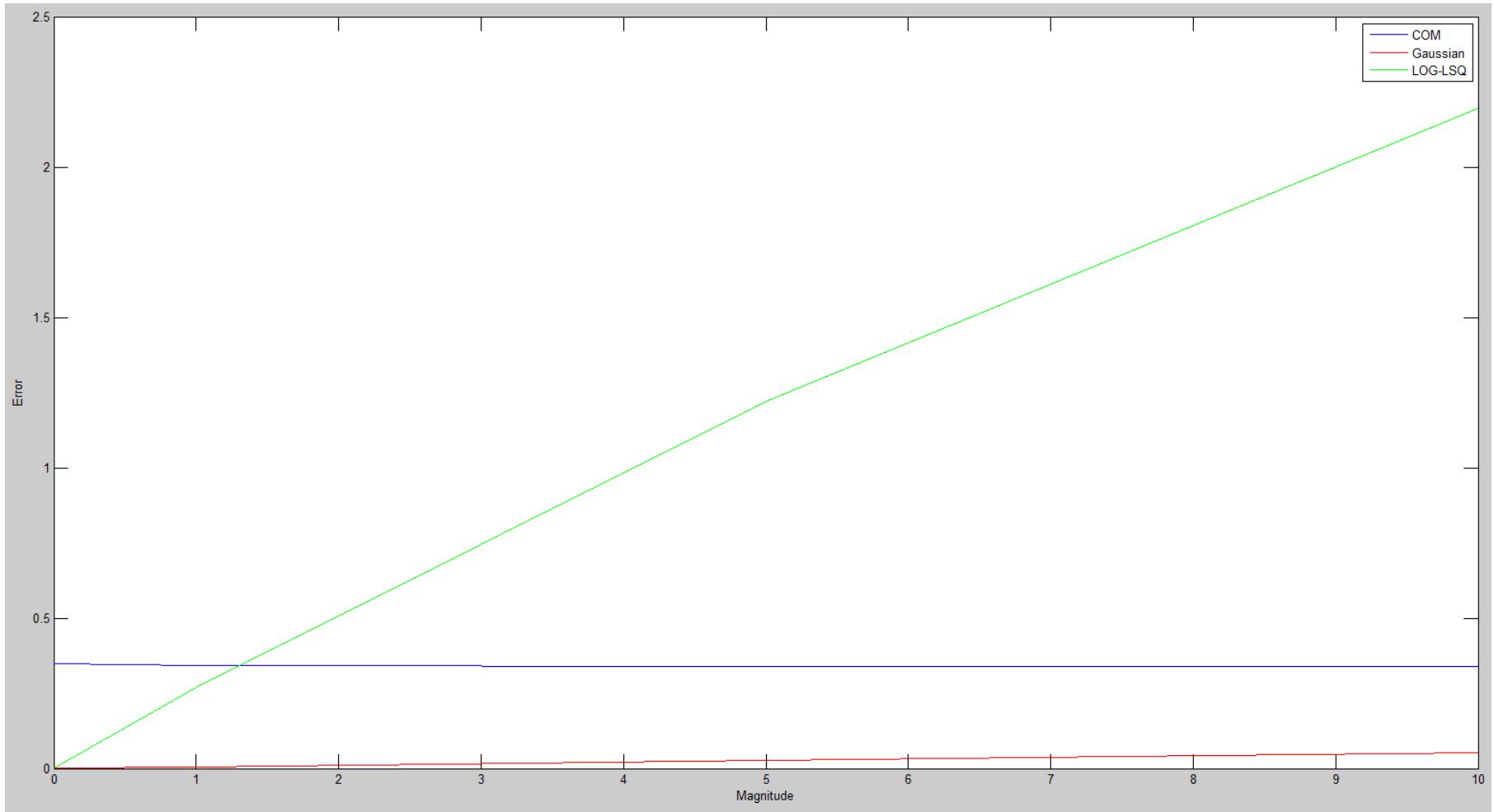
Simulation

- Mask Error (Noise 1%, m 3)



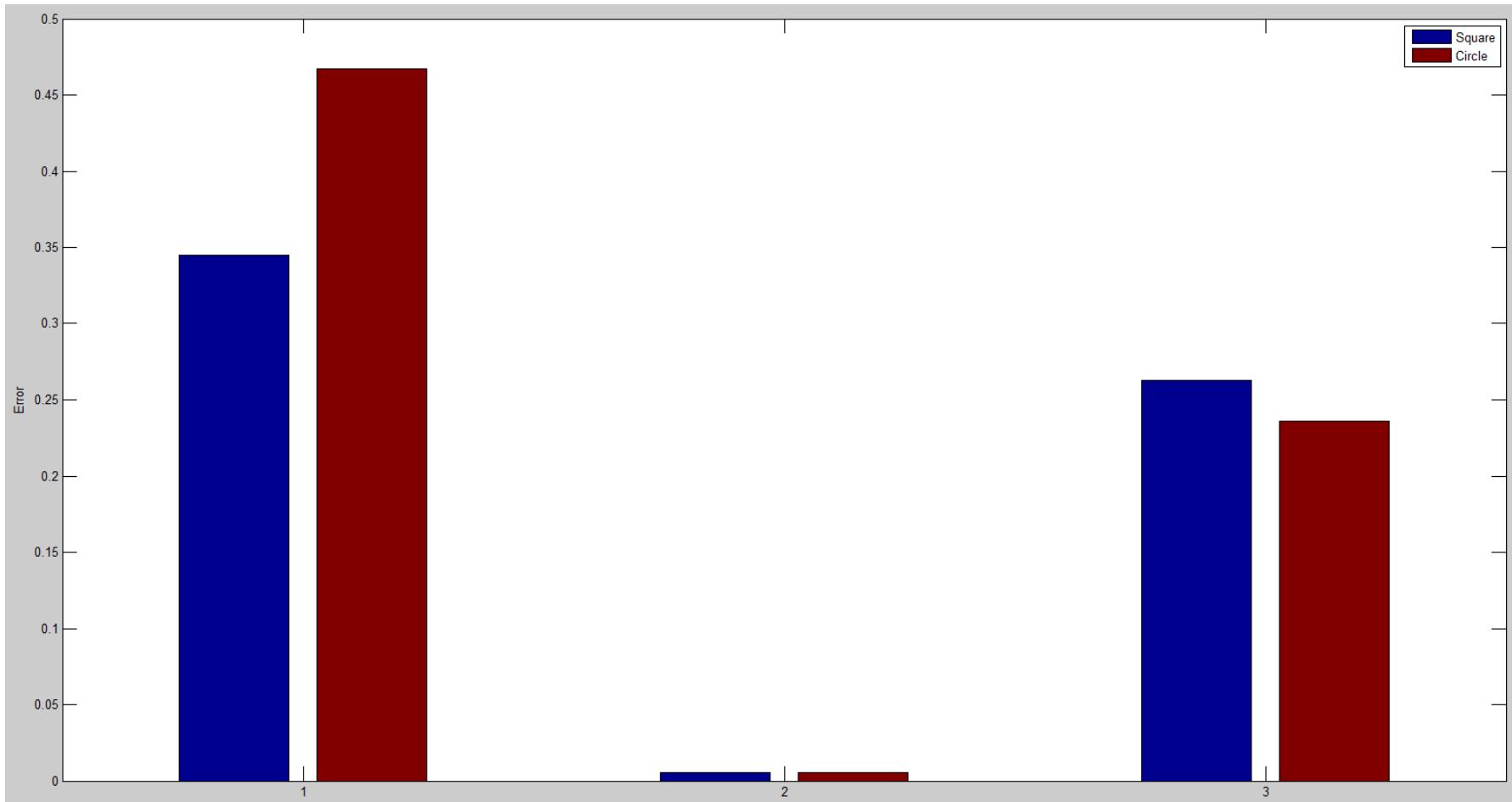
Simulation

- Noise Error (m 3, Mask 7)



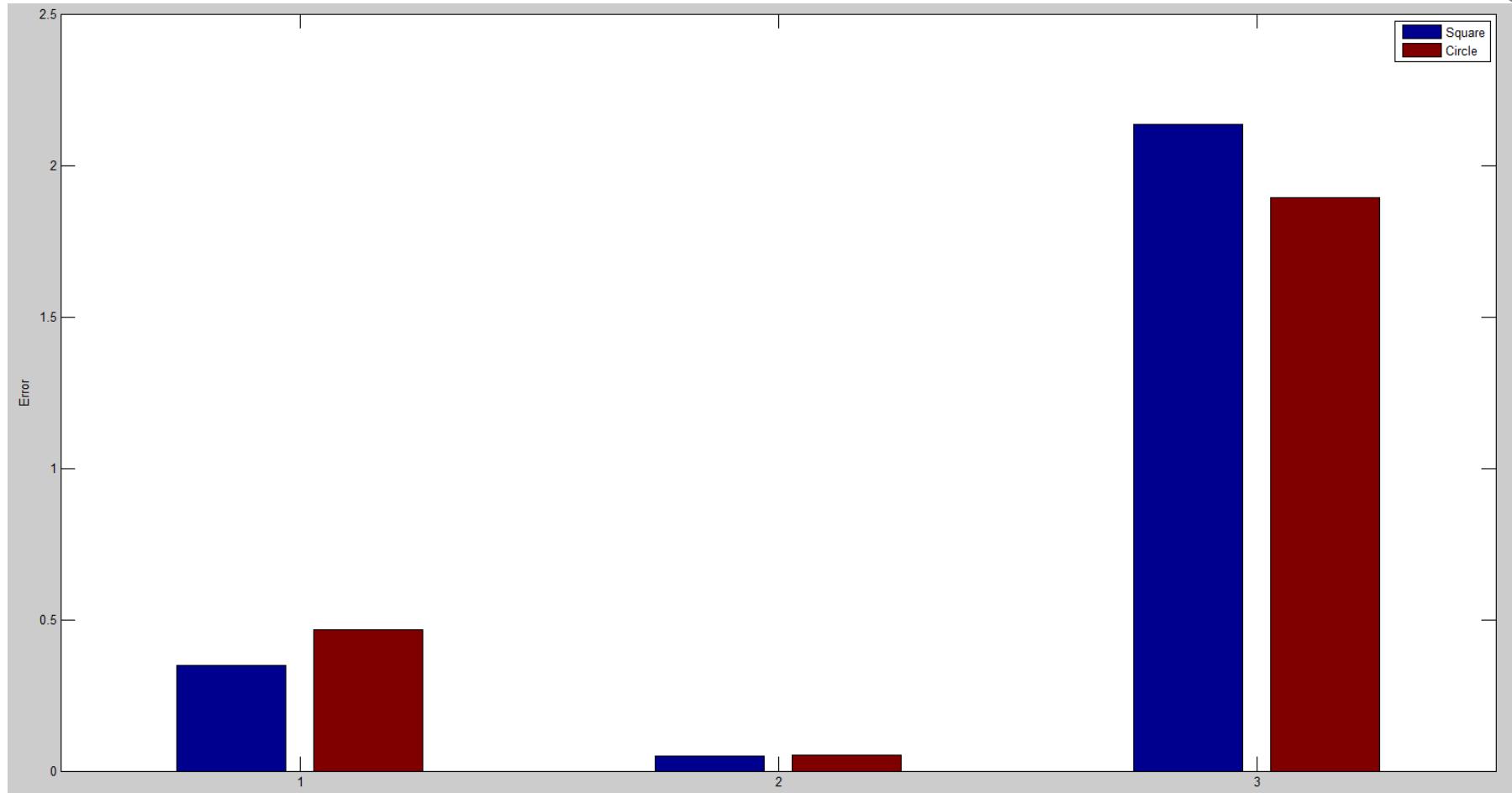
Simulation

- Mask shape (Noise 1%, m 3, Mask 7)



Simulation

- Mask shape (Noise 10%, m 3, Mask 7)



CONCOLUSION

- 수치해석 Method가 가장 정확
- COM은 어느 상황에서도 준수한 성능을 보유
- 방법론마다 Mask shape의 영향이 다름
- Mask size는 유동적으로 계산하는 것이 용이
- Sigma의 값에 따른 마스크 크기 선정